

Increasing the resolution of the FFT for more precise FLR-frequency determination in VLOS

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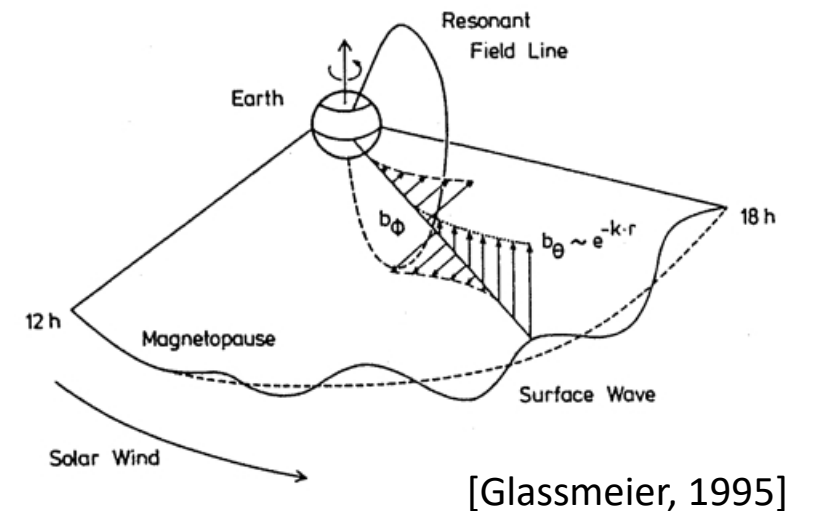
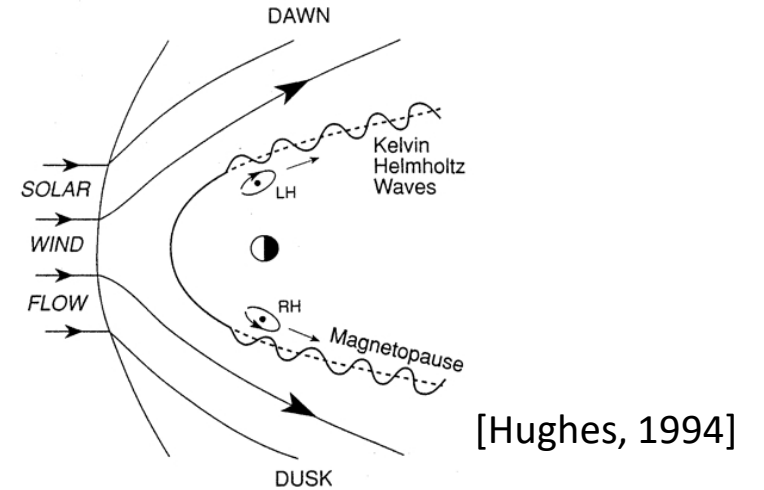
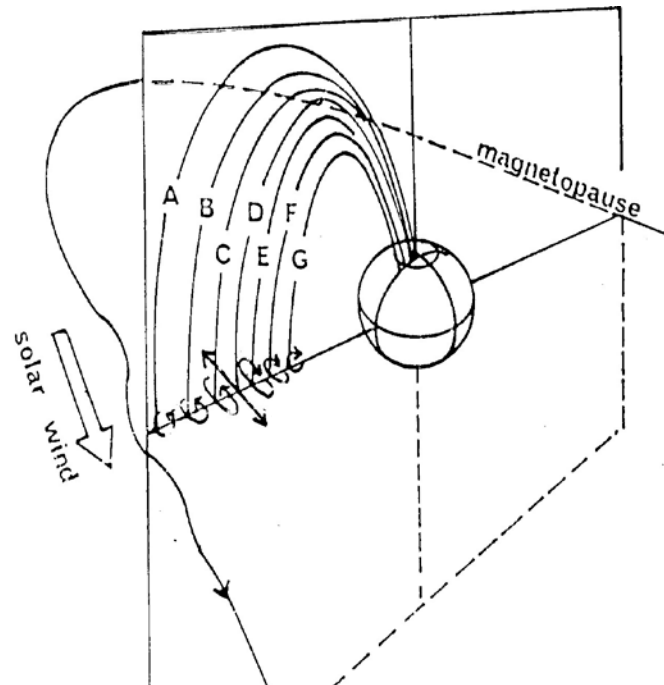
Field-Line Resonance (FLR)

- Frequency of the field-line eigen-oscillation
- Magnetospheric Density

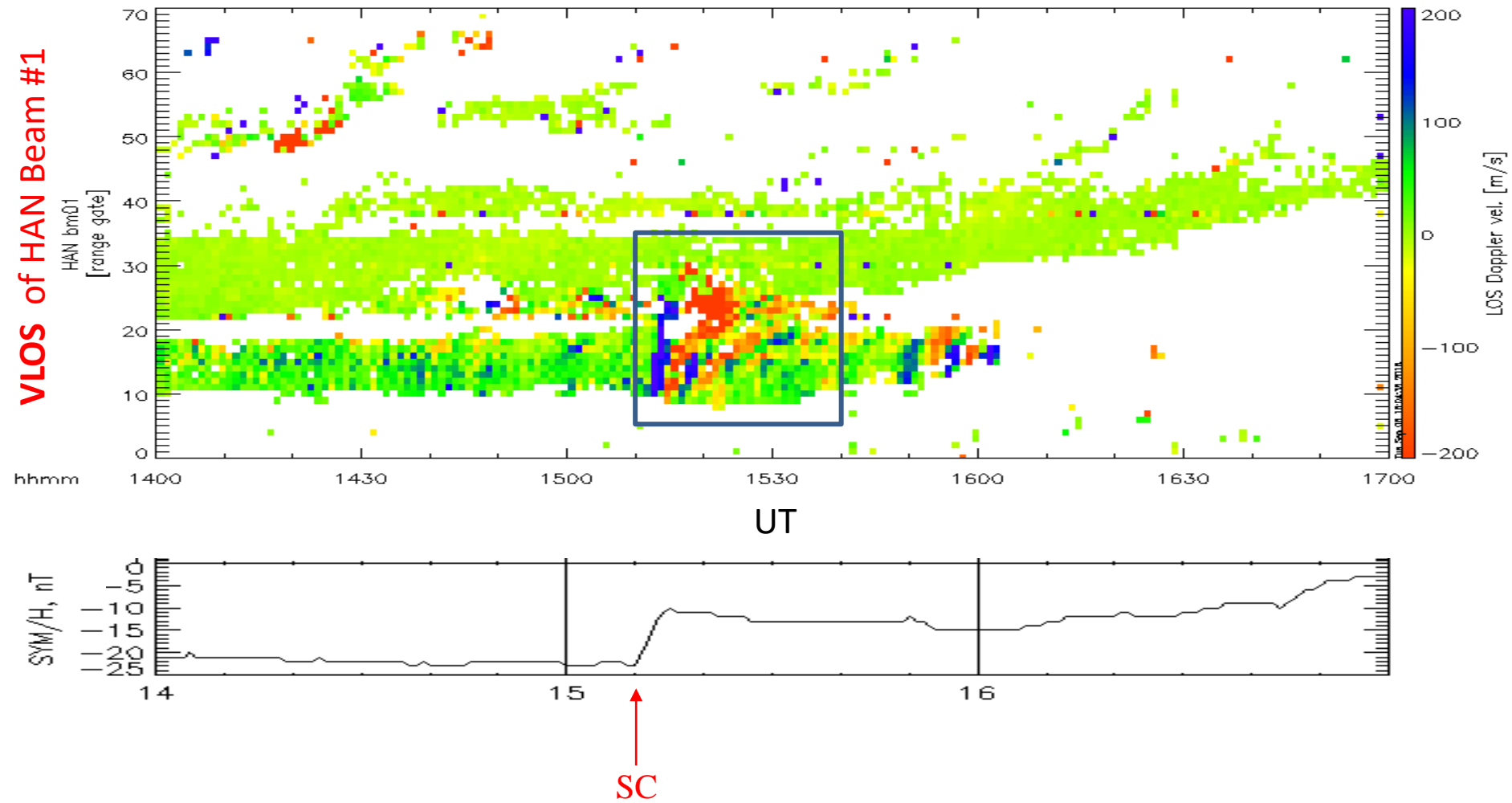
An incoming wave and a magnetospheric field-line eigen-oscillation resonates where the field line's **eigen-period** matches the **incoming-wave period**.

$$\text{Eigen-period} \propto l, \rho^{1/2},$$

where l : field-line length,
 $\rho^{1/2}$: plasma density



Event of this paper



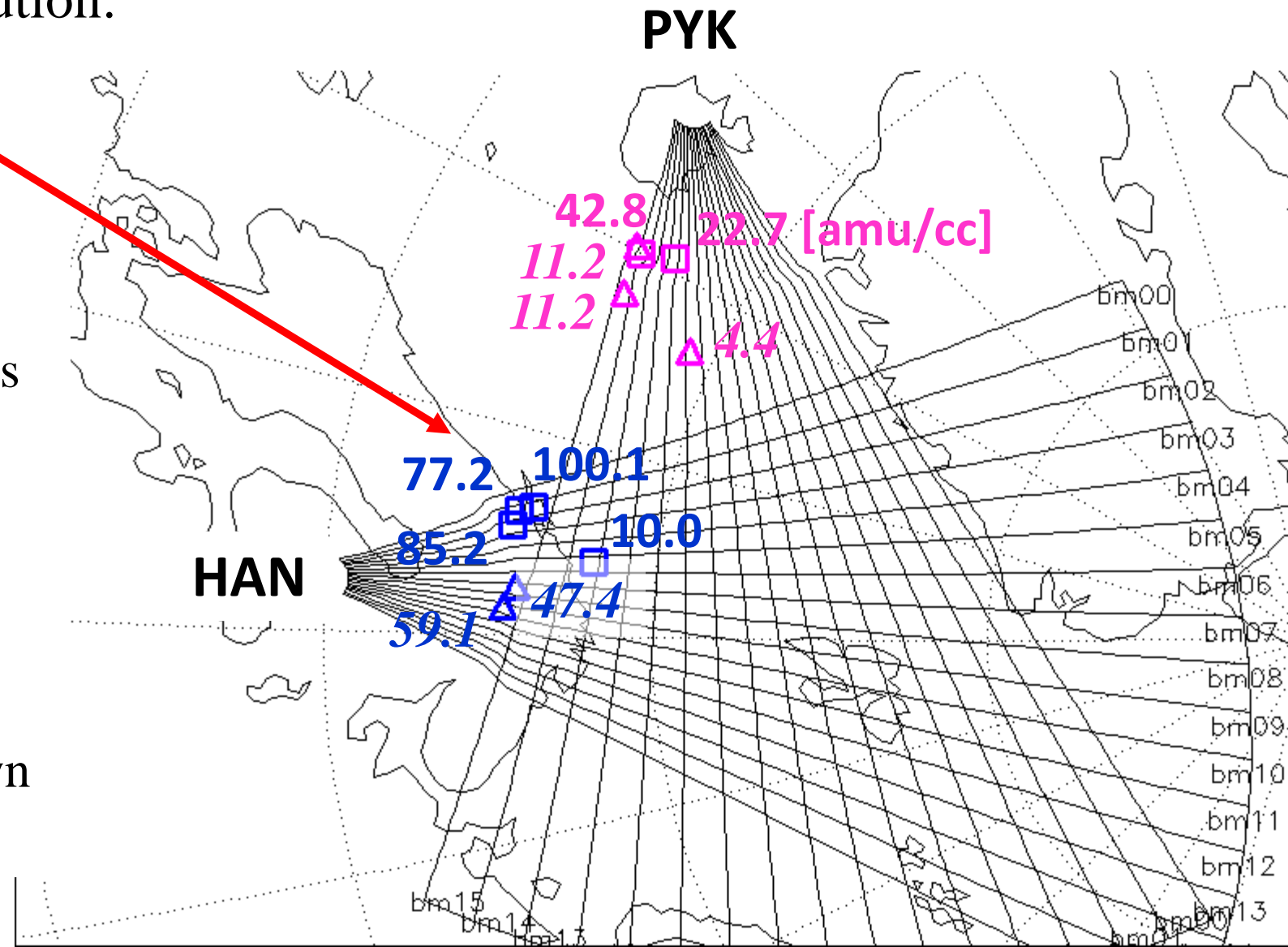
○ Density estimates from the auto-ID'ed FLR's

○ Features of the density distribution:

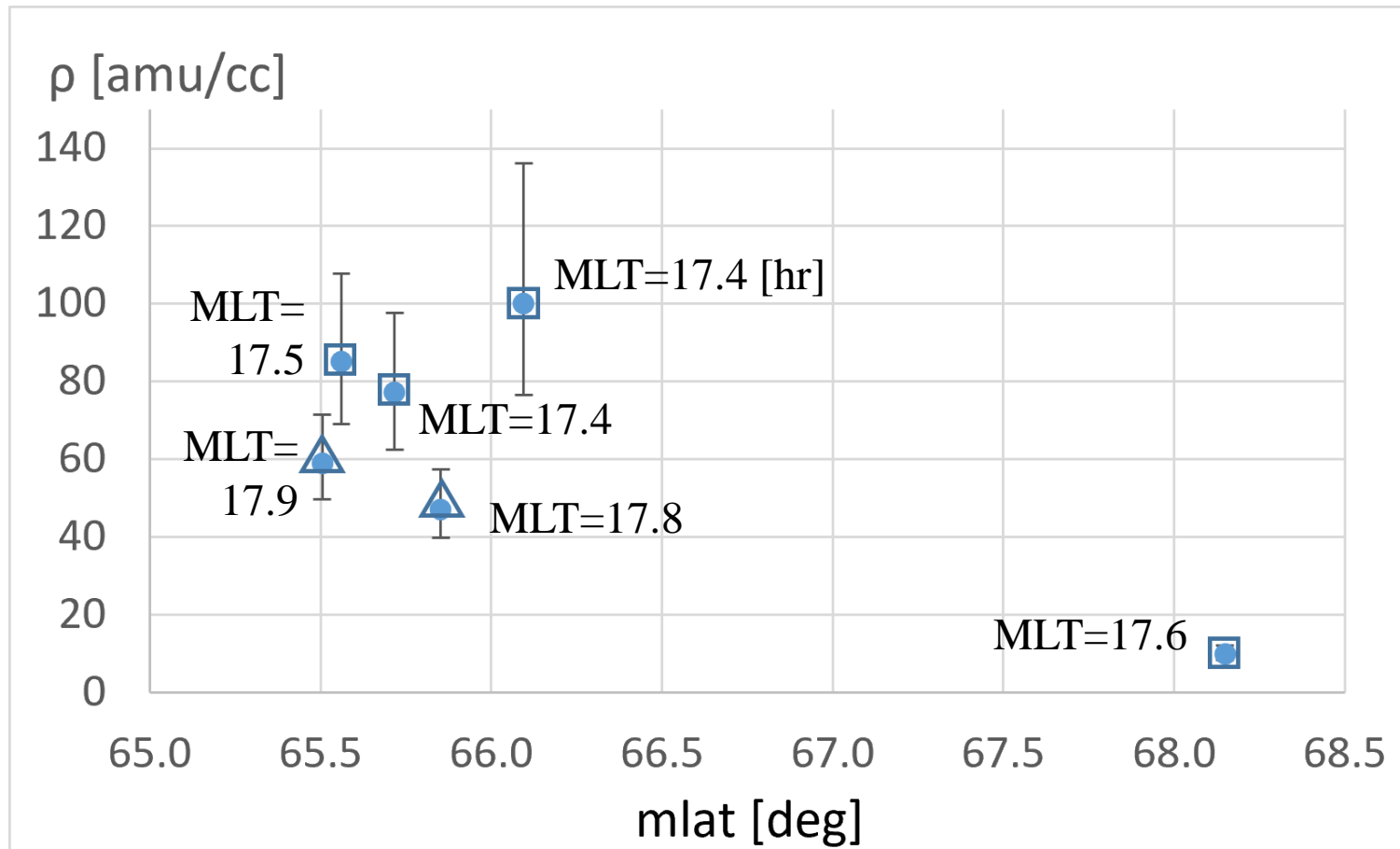
- There are pretty large density fluctuation in the small area.
- We notice that the densities estimated from the sea-backscattered signals (triangles in the figure) are smaller than those estimated from the ionosphere-backscattered signals.

→ We have estimated the error ranges of the densities shown right (→ next slide):

Backscattered from:
Square: ionosphere ($|V_{LOS}| > 190$ m/s)
Triangle: sea ($|V_{LOS}| < 35$ m/s)



- The density estimates of the last slide, shown in a different format, with largest-estimated error bars (assuming that the error in the frequency is 0.5556 / 2.0 mHz)



- The difference in the small area in the last slide (77.2 and 100.1) is within the (largest-estimated) error range.
- The difference between the ionospheric and sea backscatters are also within the error range, but the overlapping of the error bars is not so significant.

○ Density fluctuation in the small area

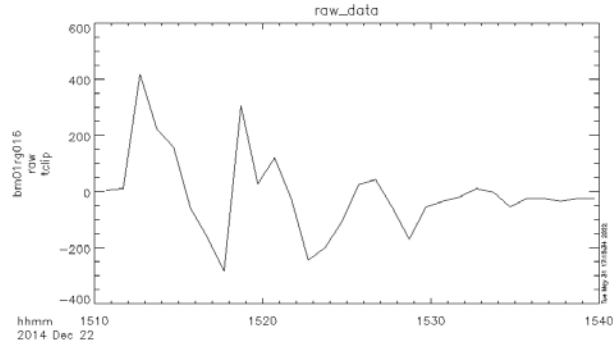
- The pretty large density fluctuation in the small area, as shown in the last two slides, could be due the quantization of the frequency ($n \times 0.5556$ mHz), coming from the FFT.

→ Needed: ▪ Error estimation (rough estimation was made in the last slide)

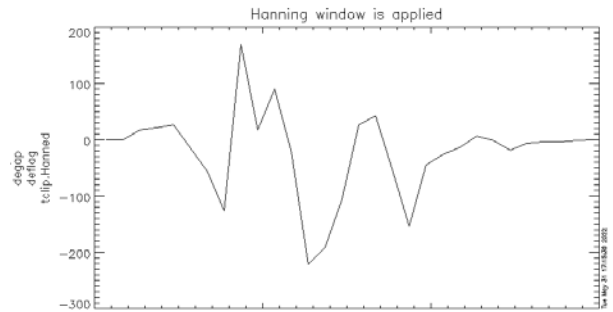
- Increasing the frequency resolution by using
 - DFT (direct Fourier Transformation) or
 - Zero-padding

→ We have tested the zero-padding (next slide):

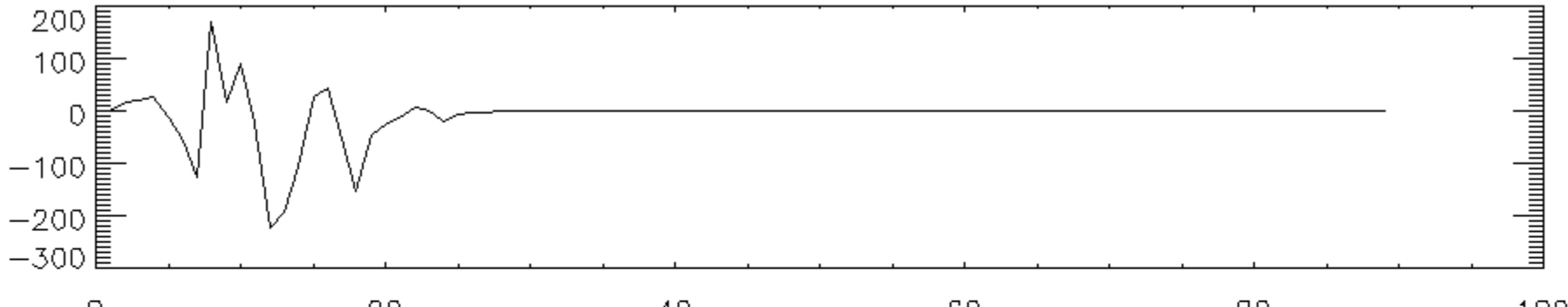
○ Zero-padding



Raw VLOS data
(Beam #01, RG #16,
15:10-15:40 UT



Hanning window
is applied



Zero-padded

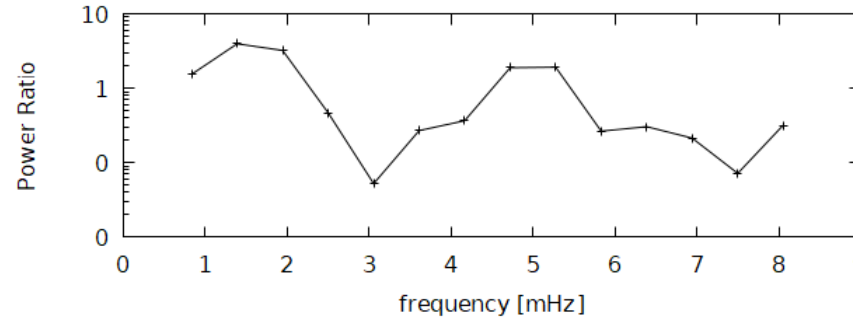
○ Zero-padding

The result of applying the **gradient method** to

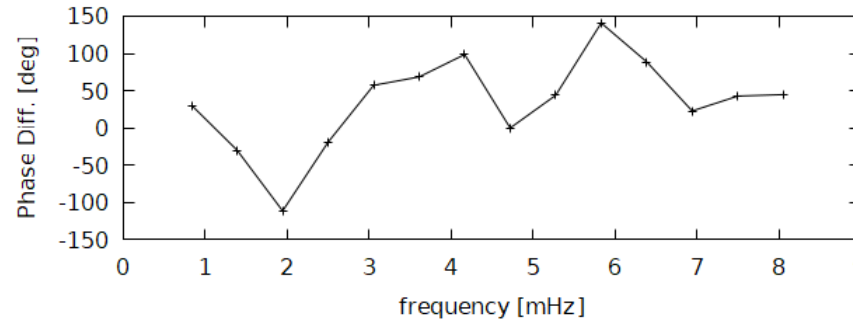
(1) two VOS timeseries-datasets from RGs #18 and #16, after the Hanning window is applied to the each.

(2) two timeseries-datasets of padding zeros to the datasets of (1)

(1)

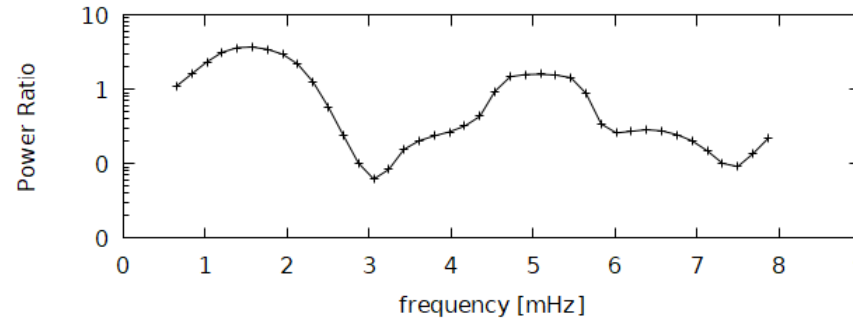


Amplitude ratio

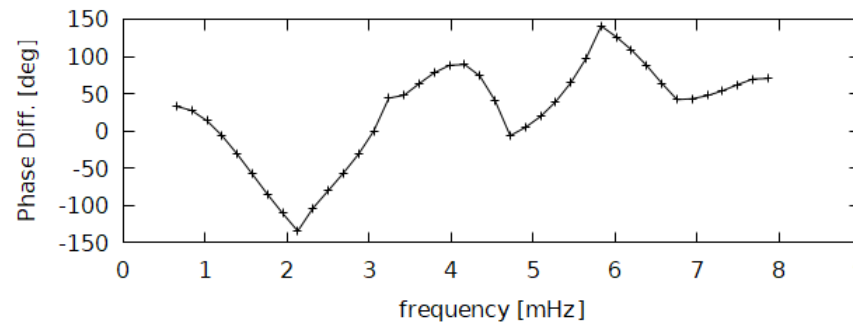


Phase difference

(2)

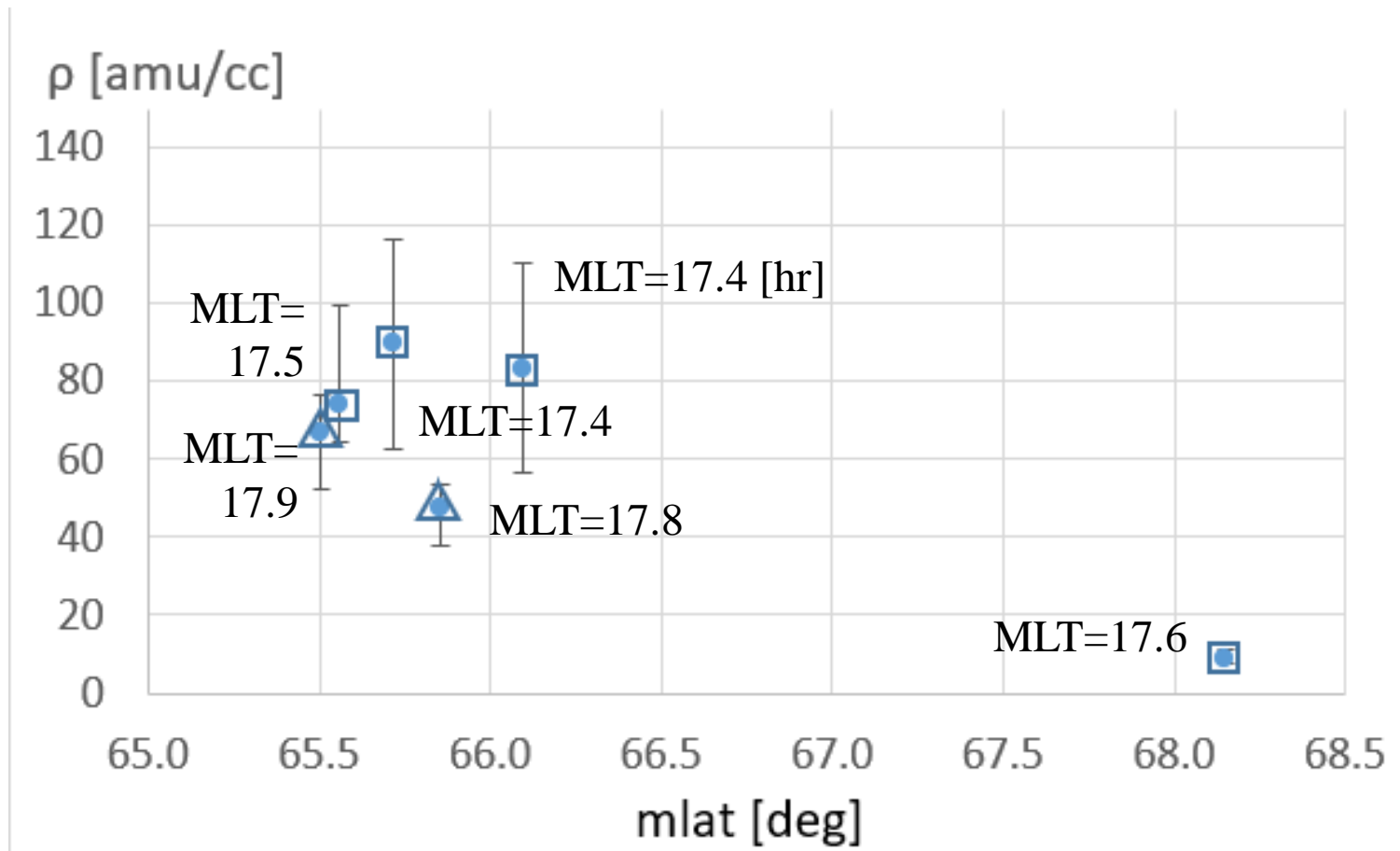


Amplitude ratio



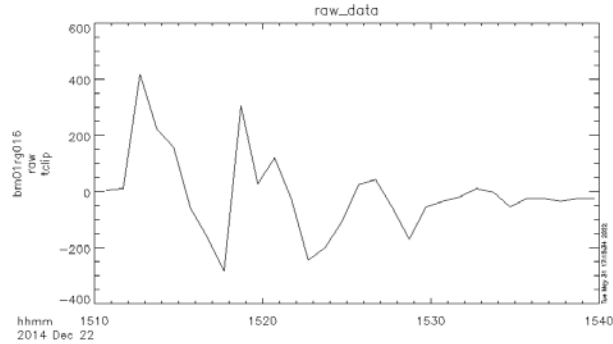
Phase difference

- The zero-padding-based density estimates, with largest-estimated error bars (assuming that the error range in the frequency is half the frequency distance between the maximum and minimum amplitude-ratio points).

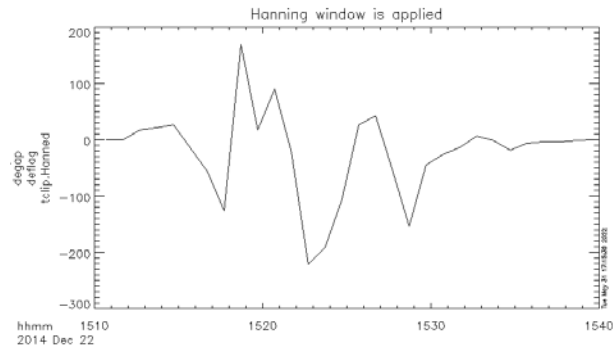


- The difference in the small area has become smaller.
- The difference between the ionospheric and sea backscatters has also become smaller.

○ Effect of applying the Hanning window

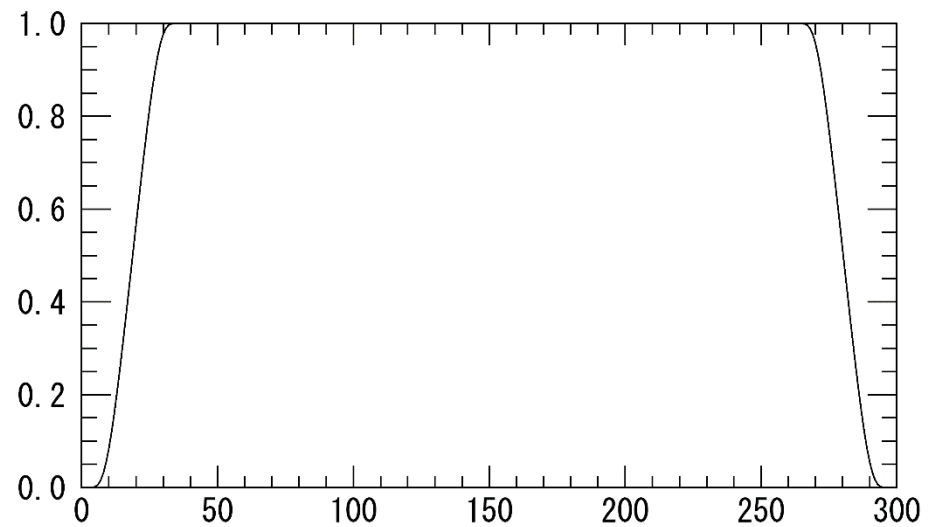


Raw VLOS data
(Beam #01, RG #16,
15:10-15:40 UT

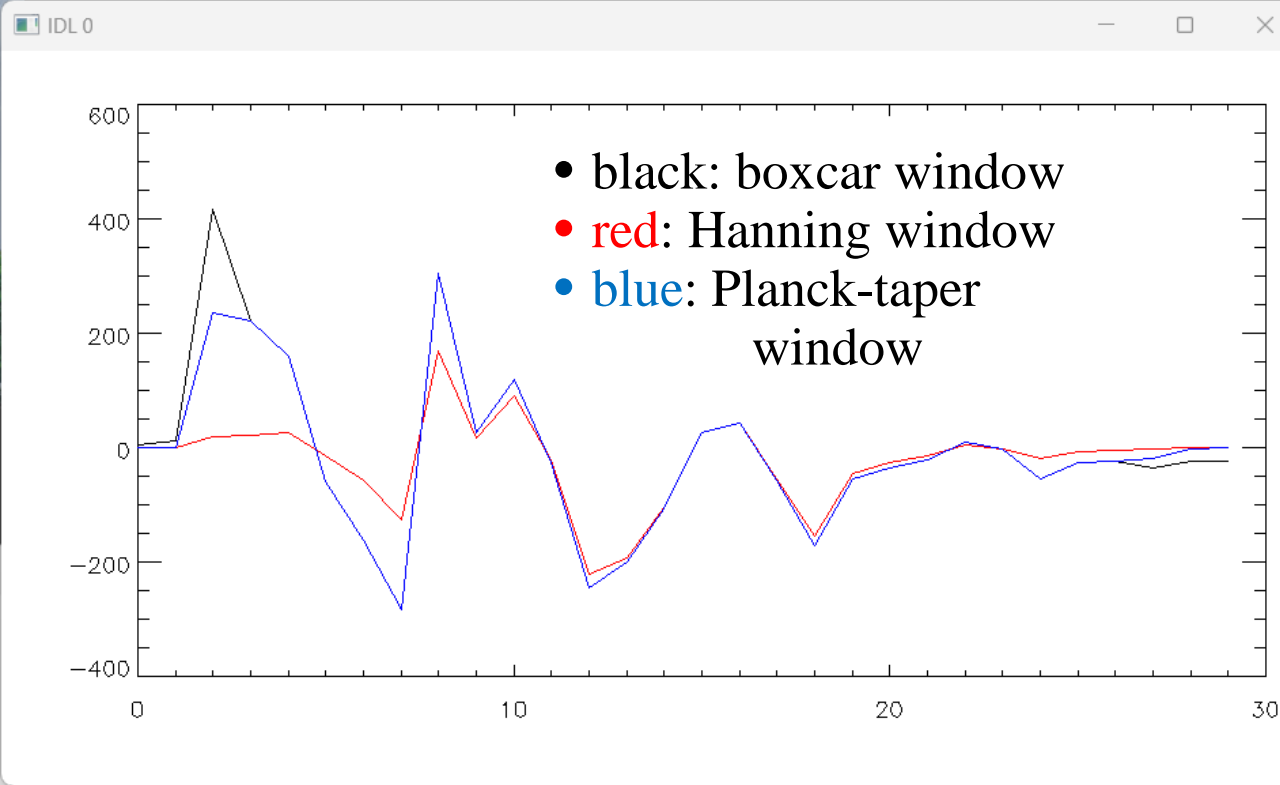


Hanning window
is applied

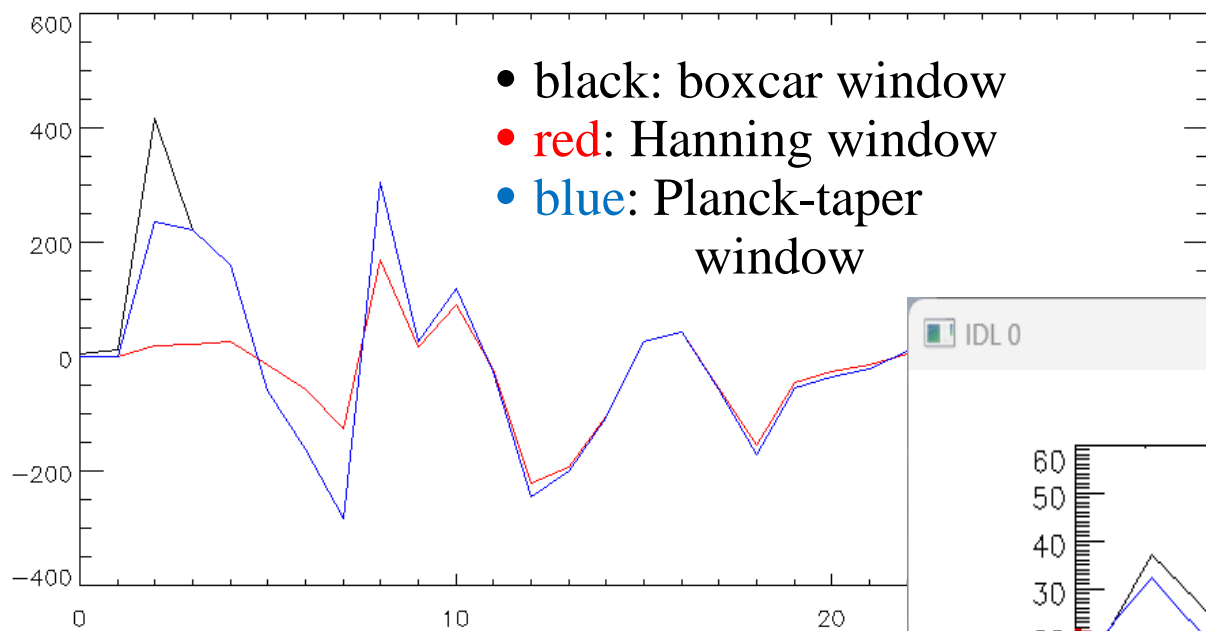
- The initial pulse just after the SC is almost deleted by the Hanning Window.
Does not this affect the FFT result?



- We have applied the Planck-taper window, and looked at the difference between the two windows

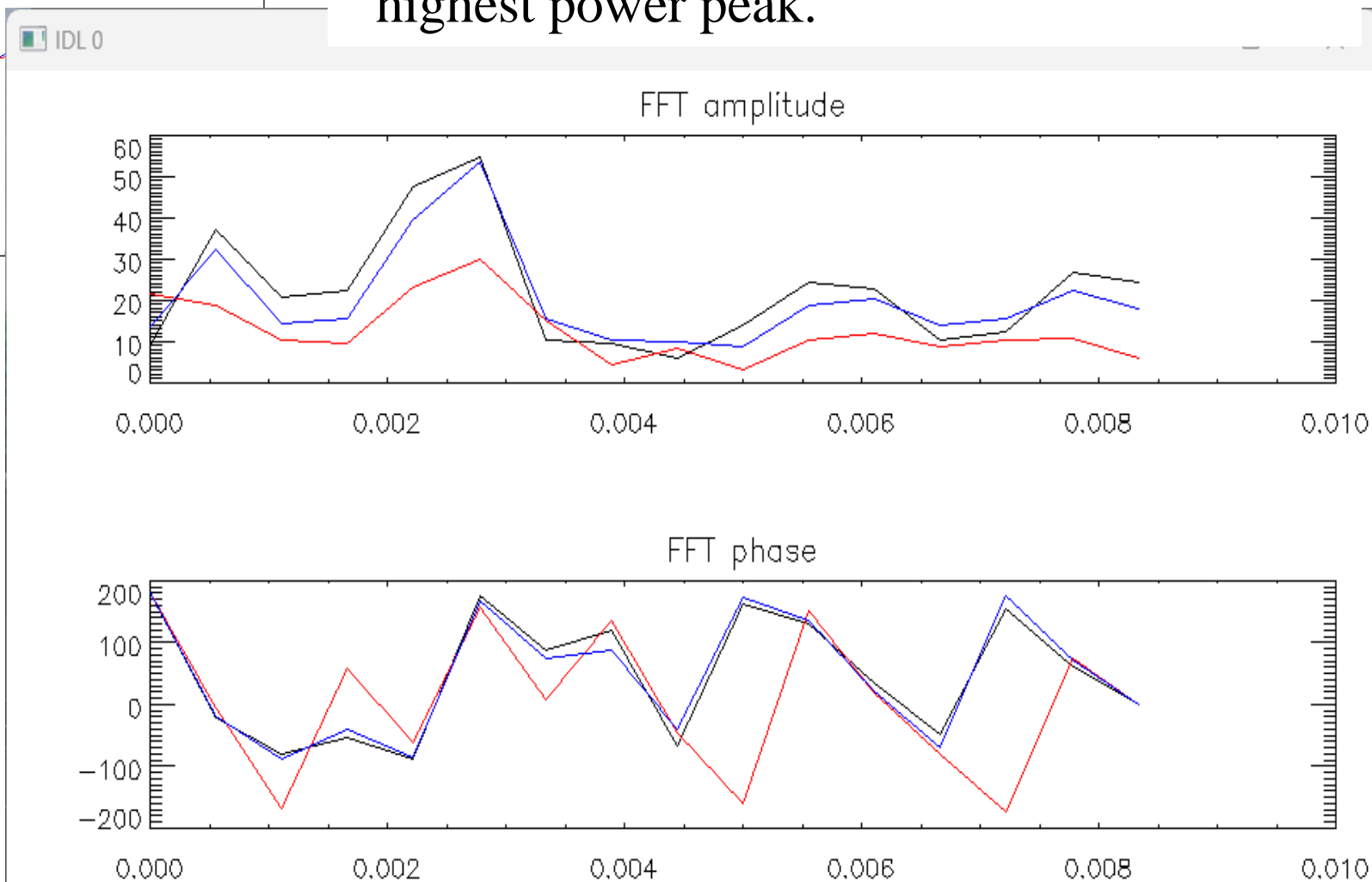


- The initial pulse is pretty well kept with the Planck-taper window, while
- it is almost deleted by the Hanning window.



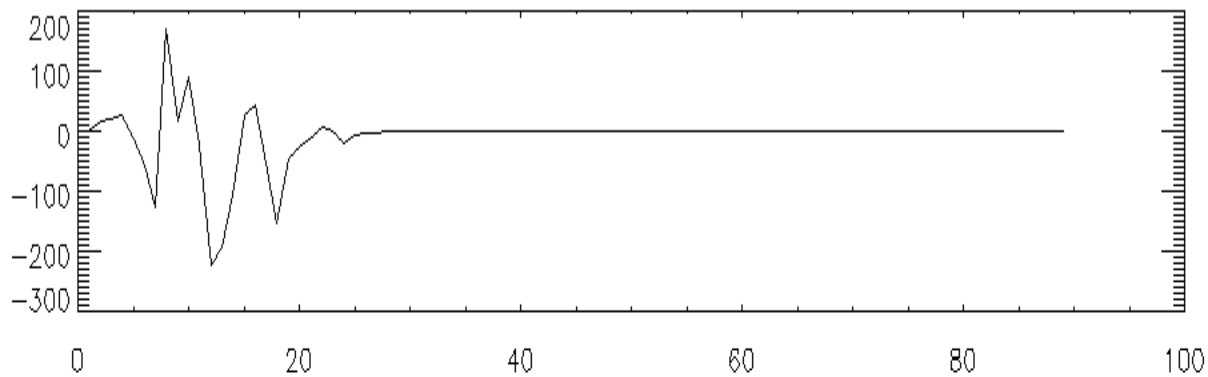
- The initial pulse is pretty well kept with the Planck-taper window, while
- it is almost deleted by the Hanning window.

- The amplitude of the highest peak is much smaller with the Hanning than the Planck-taper.
- The phase is the most different near the highest power peak.

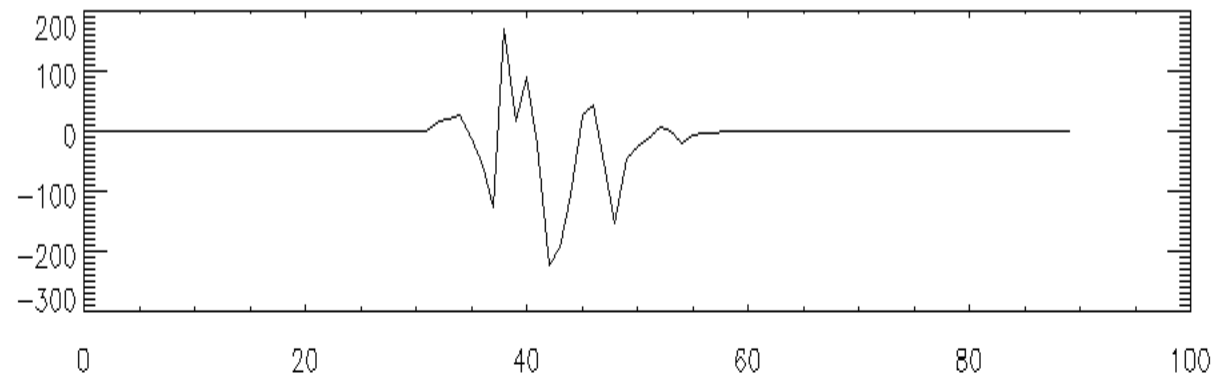


- Zero-padding before and after the true data

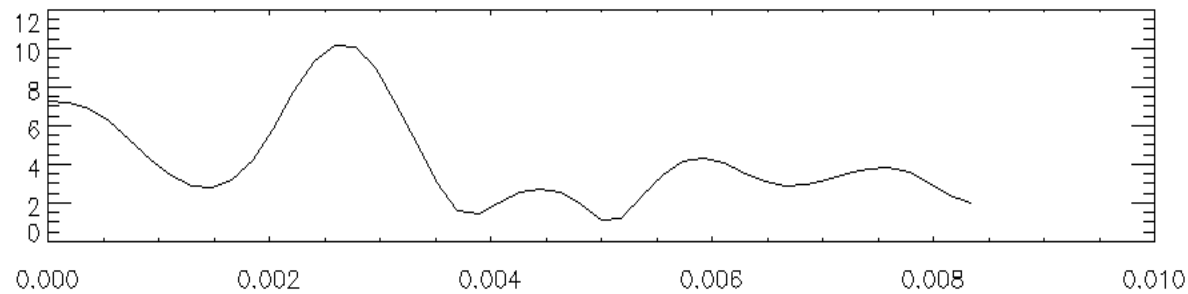
Hanning + Zero-padded



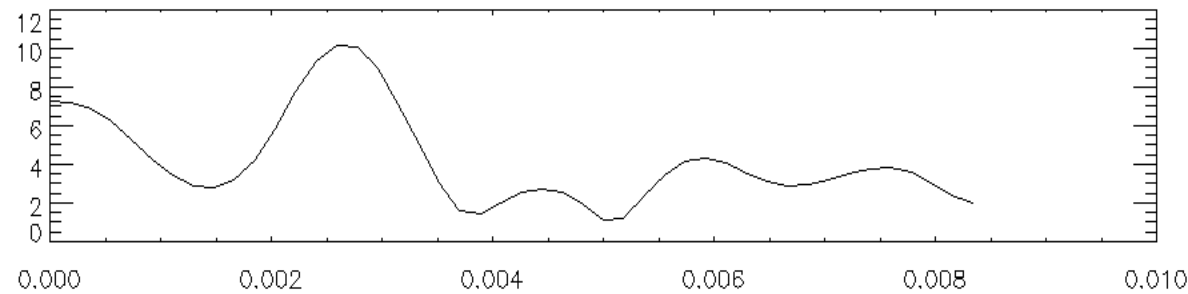
Hanning + Zero-padded



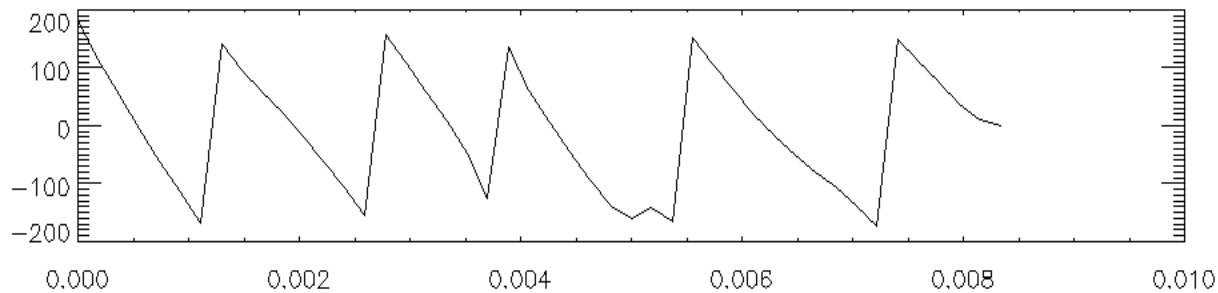
FFT of the above timeseries data, amplitude



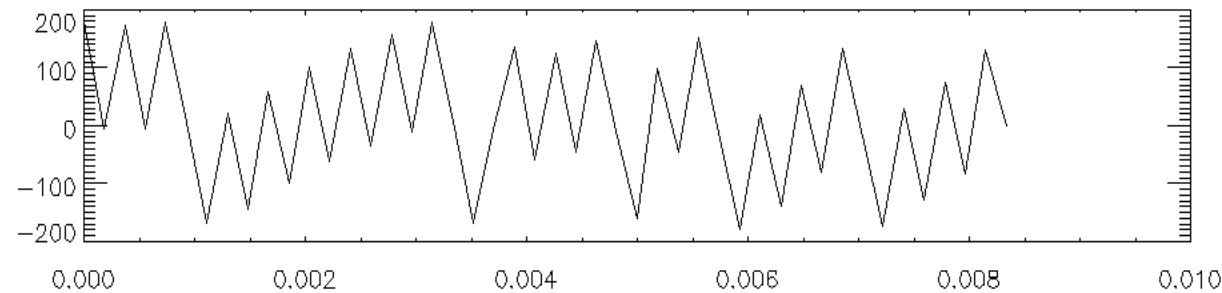
FFT of the above timeseries data, amplitude



FFT of the above timeseries data, phase



FFT of the above timeseries data, phase



○ Summary

- The fluctuations of the estimated plasma density in the small area observed by HAN have become smaller by increasing the frequency resolution.
- For SC-driven perturbations, the Hanning window could be inadequate.

○ Ongoing project

- We have been developing an all-in-one IDL code to automatically identify FLR events for all the beams of a SuperDARN radar at any given time.
- The current codes could be further improved, and we are working on that, too.

-END-